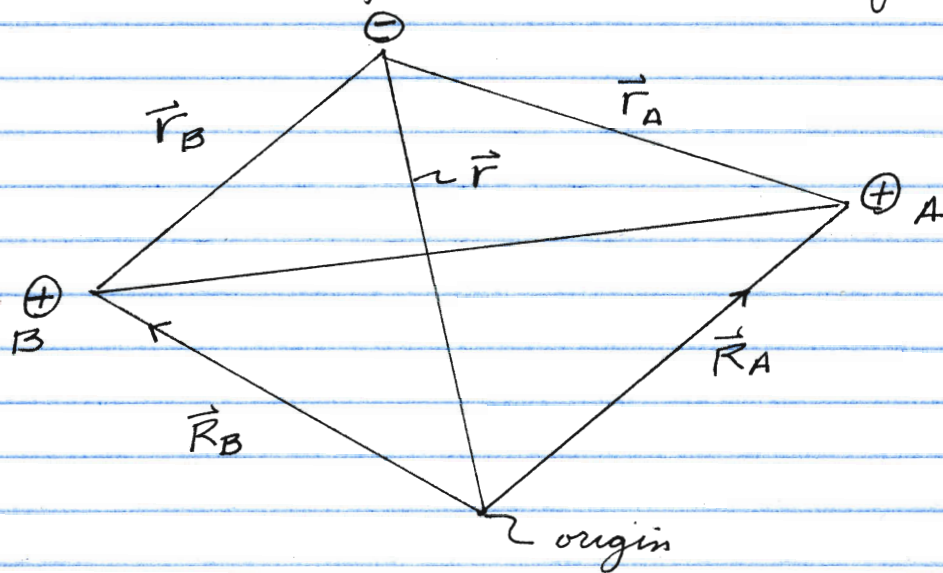


Born Oppenheimer Approximation

Consider the complete Hamiltonian for H_2^+



The separation between A & B is $R = |\vec{R}_A - \vec{R}_B|$

$$\hat{H}_T = -\frac{1}{2m_p} \nabla_A^2 - \frac{1}{2m_p} \nabla_B^2 - \frac{1}{2} \nabla^2 - \frac{1}{r_A} - \frac{1}{r_B} + \frac{1}{R}$$

~~The with~~ The Born Oppenheimer approximation recognizes that the mass of the proton m_p is much larger than the mass of the electron & as a result we may consider the nuclei ^{as} stationary & solve for the motion of the electron for a fixed R .

The Born-Oppenheimer Hamiltonian for H_2^+ is

$$\hat{H}_{B0} = -\frac{1}{2} \nabla^2 - \frac{1}{r_A} - \frac{1}{r_B} + \frac{1}{R}$$

If we solve

$$\hat{H}_{B0} \psi_{B0} = E_{B0} \psi_{B0}$$

we must first fix R & then solve the equation.

$$\text{Let } \hat{H}_e = -\frac{1}{2} \nabla^2 - \frac{1}{r_A} - \frac{1}{r_B}$$

$$\text{Then } \hat{H}_{B0} = \hat{H}_e + \frac{1}{R}$$

So

$$(\hat{H}_e + \frac{1}{R}) \psi_{B0} = E_{B0} \psi_{B0}$$

$$\hat{H}_e \psi_{B0} = (E_{B0} - \frac{1}{R}) \psi_{B0} = E_e \psi_{B0}$$

Since

$$r_A = \sqrt{(x-x_A)^2 + (y-y_A)^2 + (z-z_A)^2}$$

$\hat{H}_e$, ψ_{B0} & E_{B0} all depend on R , so

$$E_{B0}(R) = E_e(R) + \frac{1}{R}$$

We know $E_e(R)$ at two values of R .

First, when $R = \infty$ $\epsilon_e = -\frac{1}{2}$

because when H_2^+ dissociates it forms $H \cdot + H^+$
 & the energy of $H \cdot$ is $-\frac{1}{2}$.

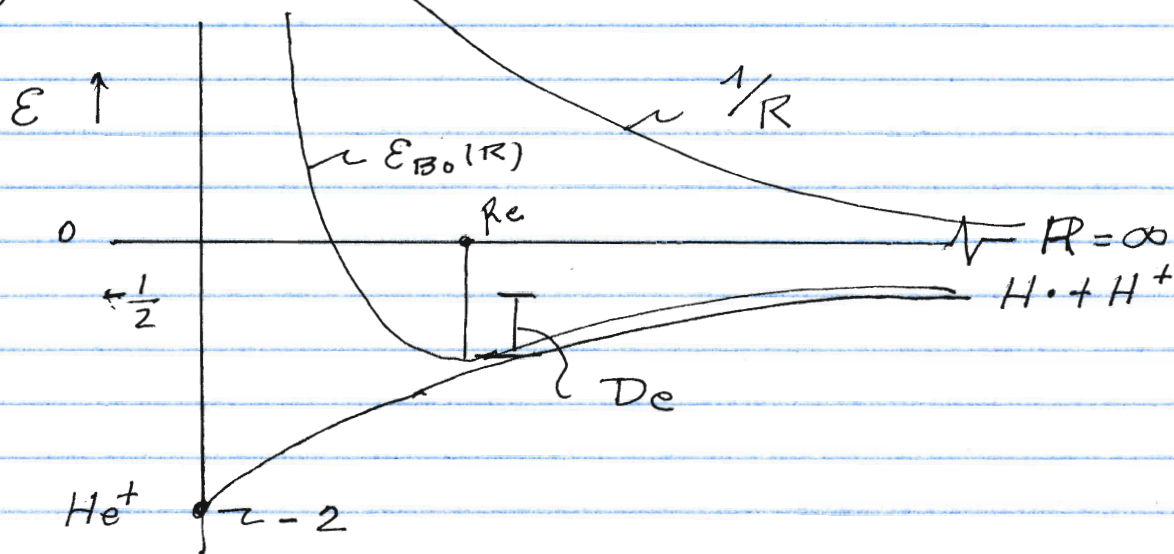
Second, when $R=0$. In the case $r_A = r_B$

& $\hat{H}_{He} = -\frac{1}{2} \nabla^2 - \frac{2}{r_A}$ & this is the He^+

Hamiltonian. Consequently $\epsilon_e(0) = -\frac{1}{2} \cdot 4 = -2$

The $R = \infty$ solution is called the separated atom limit while the $R=0$ is the united atom limit.

The eigenvalue $\epsilon_e(R)$ varies smoothly between these limits. The total E_{B_0} is then the sum of the negative ϵ_e & the positive $1/R$



Once one has $E_{B0}(R)$ one forms the Hamiltonian for the nuclear motion

$$\hat{H}_N = -\frac{1}{2M_P} \nabla_A^2 - \frac{1}{2M_P} \nabla_B^2 + E_{B0}(R)$$

§ Solves

$$\hat{H}_N \psi(R) = E(R) \psi(R)$$

Once can write $\hat{H}_N$ as $\hat{H}_{cm} + \hat{H}_{ROT} + \hat{H}_{VIB}$

§

$$\psi(R) = \psi_{cm} \psi_{ROT} \psi_{VIB}$$

$$E(R) = E(R_0) + E_{cm} + E_{ROT} + E_{VIB}$$

$$E(R_0) = E_{B0}(R_0)$$

E_{cm} = kinetic energy of center of mass

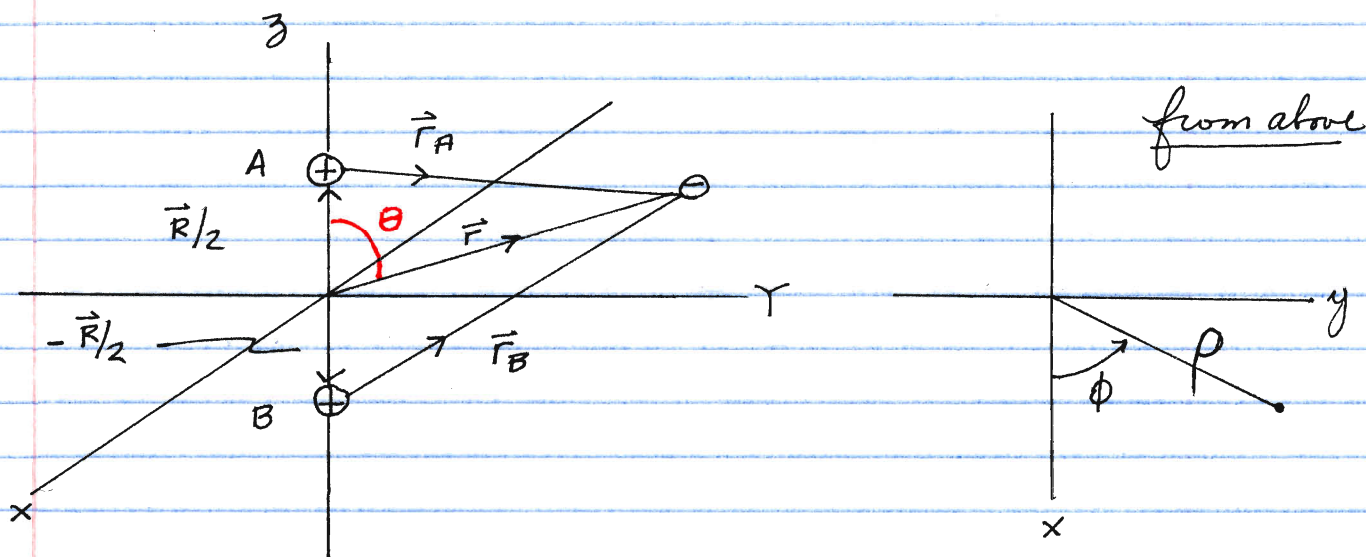
$$E_{ROT} = \frac{2J(J+1)}{2I} \frac{\hbar^2}{2I} \text{ as before}$$

$$E_{VIB} = \hbar \nu (v + \frac{1}{2}) + \dots \text{ as before}$$

We will focus on $\hat{H}_{B0}$ & simply call it $\hat{H}$.

Lets see what we can learn about the solutions to $\hat{H}\psi = E\psi$ in the BO approximation

Lets place the protons on the z axis with one at $z = +R/2$ & the other at $z = -R/2$



The x, y, z coordinates of the electron can be expressed in cylindrical coordinates

$$x = \rho \cos \phi, \quad y = \rho \sin \phi \quad \& \quad z = z$$

& ρ_0

$$-\frac{1}{2} \nabla^2 = -\frac{1}{2\rho^2} \frac{\partial}{\partial \rho} \left(\rho^2 \frac{\partial}{\partial \rho} \right) + \frac{1}{\rho^2} \frac{\partial^2}{\partial \phi^2} + \frac{\partial^2}{\partial z^2}$$

recall $\hat{L}_z = -i \frac{\partial}{\partial \phi}$ in atomic units

$$\rho \quad \hat{L}_z^2 = -\frac{\partial^2}{\partial \phi^2}$$

& since $-\frac{1}{2} \nabla^2$ contains $\frac{\partial^2}{\partial \phi^2}$ the two operators

commute. What about $\frac{1}{r_A} + \frac{1}{r_B}$

$$r_A = \sqrt{x^2 + y^2 + (z - R/2)^2} = \sqrt{\rho^2 + (z - R/2)^2}$$

&

r_A & r_B do not depend on ϕ so

$$[\hat{L}_z, \hat{L}_z] = 0$$

This means

$$\psi(\rho, z, \phi) = \psi(\rho, z) e^{iM\phi}$$

with $M = 0, \pm 1, \pm 2, \dots$ & the eigenfunctions

must be eigenfunction of L_z with eigenvalue M .

M	0	± 1	± 2	± 3	$\dots$
letter	σ	π	δ	ϕ	$\dots$
α	Σ	Π	Δ	Φ	$\dots$

Consider the consequence of the two nuclei being identical.

Define an inversion operator $\hat{I}$ which is such that

$$\hat{I} f(\vec{r}) = f(-\vec{r})$$

The eigenvalue problem for $\hat{I}$ is

$$\hat{I} \chi(\vec{r}) = \lambda \chi(\vec{r})$$

but $\hat{I} \chi(\vec{r}) = \chi(-\vec{r})$ by definition. & so

$$\hat{I}^2 \chi(\vec{r}) = \hat{I} \chi(-\vec{r}) = \chi(\vec{r})$$

but $\hat{I}^2 \chi(\vec{r}) = \hat{I} \lambda \chi(\vec{r}) = \lambda^2 \chi(\vec{r})$

so $\lambda^2 = 1 \quad \& \quad \lambda = \pm 1$

Consider now the commutator of $\hat{I}$ & $\hat{H}$.

$\hat{I}$ certainly commutes with ∇^2 because of the second derivatives $\frac{\partial^2}{\partial x^2}$ etc. What about V_A & V_B .

$$\begin{aligned} \hat{I} V_A &= \hat{I} \sqrt{x^2 + y^2 + (z - R/2)^2} \\ &= \sqrt{x^2 + y^2 + (-z - R/2)^2} = V_B \end{aligned}$$

& conversely $\hat{I} V_B = V_A$

So the operator $\hat{I}$ commutes with the Hamiltonian for H_2^+ . ~~There is~~ Then if

$$\hat{H}\psi = E\psi$$

we have some ψ which change sign

$$\hat{I}\psi = -\psi$$

& some that don't

$$\hat{I}\psi = +\psi$$

Eigenfunctions that change sign upon inversion are called ungerade (u) functions while those that do not change sign are gerade or g functions.

It is reasonable to assume that the lowest energy state of H_2^+ starts out as the 1s orbital at $R=\infty$ and finishes as the 1s orbital of He^+ at $R=0$. The 1s of He^+ has g symmetry & the 1s of H has $m=0$ & is a σ orbital. The electronic configuration is then $1\sigma_g^1$.

Just as with atoms, diatomic molecules can have term symbols. For H_2^+ we have one electron & so $2S+1=2$, a doublet. The ground state term is $^2\Sigma_g^+$ where the + means the wavefunction does not change sign when reflected through any plane containing the two nuclei.

Approximate Wave functions for H_2^+

If we take our variation functions as

$$\psi = C_A 1A_A + C_B 1A_B$$

where

$$1A_A = \frac{1}{\sqrt{\pi}} e^{-r_A} \quad \& \quad 1A_B = \frac{1}{\sqrt{\pi}} e^{-r_B}$$

we can determine C_A & C_B in two ways. First, we can insist that ψ be an eigenfunction of $\hat{I}$ & second we can use the variation method.

Eigenfunctions of $\hat{I}$

recall

$$\hat{I} r_A = r_B \quad \& \quad \hat{I} r_B = r_A$$

$$\text{So } \hat{I} 1A_A = 1A_B \quad \& \quad \hat{I} 1A_B = 1A_A$$

So

$$\hat{I} \psi = C_A \hat{I} 1A_A + C_B \hat{I} 1A_B = C_A 1A_B + C_B 1A_A$$

If we require ψ to be ψ_g

$$\hat{I} \psi = +\psi$$

∴

$$C_A 1A_B + C_B 1A_A = C_A 1A_A + C_B 1A_B$$

$$(C_A - C_B) 1A_B + (C_B - C_A) 1A_A = 0$$

$$(C_A - C_B) (1A_B - 1A_A) = 0$$

for this to be true for all x, y, z , $C_A = C_B$

So $\psi_g = C_A (1A_A + 1A_B)$

we fix C_A by normalizing ψ_g

$$\langle \psi_g | \psi_g \rangle = C_A^2 \langle 1A_A + 1A_B | 1A_A + 1A_B \rangle = 1$$

So $C_A = \frac{1}{\sqrt{2+2S}}$ where $S = \langle 1A_A | 1A_B \rangle$

S is called the overlap integral & for the orbital

$1A_A$ & $1A_B$ has the form

$$S = e^{-R} \left(1 + R + \frac{R^2}{3} \right)$$

Having ψ_g we can compute the energy as a function of R . The energy is given by

$$E_g = \langle \psi_g | \hat{H} | \psi_g \rangle = \frac{\langle 1s_A + 1s_B | \hat{H} | 1s_A + 1s_B \rangle}{2 + 2S}$$

Consider the numerator

$$\langle 1s_A + 1s_B | \hat{H} | 1s_A + 1s_B \rangle = H_{AA} + H_{BB} + 2H_{AB}$$

Now

$$H_{AA} = \langle 1s_A | -\frac{1}{2} \nabla^2 - \frac{1}{r_A} - \frac{1}{r_B} + \frac{1}{R} | 1s_A \rangle$$

$$H_{AA} = -\frac{1}{2} - \langle 1s_A | \frac{1}{r_B} | 1s_A \rangle + \frac{1}{R}$$

- $\langle 1s_A | \frac{1}{r_B} | 1s_A \rangle$ = the attraction of the electron on center A
(with $\frac{1}{R}$)
for nucleus B. Sometimes called (in your text)

the Coulomb integral $J = -\langle 1s_A | \frac{1}{r_B} | 1s_A \rangle + \frac{1}{R}$

$$H_{AA} = -\frac{1}{2} + J$$

Likewise

$$H_{BB} = -\frac{1}{2} + J$$

since by symmetry $\langle 1s_B | \frac{1}{r_A} | 1s_B \rangle = \langle 1s_A | \frac{1}{r_B} | 1s_A \rangle$

$$H_{AB} = -\frac{S}{2} - \langle 1s_A | \frac{1}{r_A} | 1s_B \rangle + \frac{S}{R}$$

The quantity $-\langle 1s_A | \frac{1}{r_A} | 1s_B \rangle + S/R = K$

where K is called (in your text) the exchange integral.

The numerator is then

$$-1 + 2J - S + 2K = -(1+S) + 2(J+K)$$

Dividing by $2+2S$

$$E_g = \underbrace{-\frac{1}{2}}_{\text{energy of } H_2^+ \text{ at } R=\infty} + \underbrace{\frac{J+K}{1+S}}_{\text{energy change upon molecule formation}}$$

Define $\Delta E_+ = E_g + \frac{1}{2} = \frac{J+K}{1+S} = \Delta E_+(R)$

One can evaluate $J = e^{-2R} (1 + \frac{1}{R})$

$$K = \frac{S}{R} - e^{-R} (1+R)$$

The minimum occurs at $\frac{\partial \Delta E_+}{\partial R} = 0$

we find $R_{\min} = R_e = 2.50 a_0$ & $\Delta E_+(R_e) = -1.76 \text{ eV}$

H_2^+ is bound in this approximation but since the experimental

$R_e = 2.00 a_0$ & experimental $\Delta E_+ = -2.79 \text{ eV}$ we are not

close.

We can improve these results by generalizing the trial function & introducing a variational parameter

Let $\psi_A = \sqrt{\frac{\alpha^3}{\pi}} e^{-\alpha r_A}$ & $\psi_B = \sqrt{\frac{\alpha^3}{\pi}} e^{-\alpha r_B}$

We find $\Delta E_+(R, \alpha)$

So forming $\frac{\partial \Delta E_+}{\partial R} = 0$ & $\frac{\partial \Delta E_+}{\partial \alpha} = 0$

we find $\alpha_{opt} = 1.24$ & $R_e = 2.02 a_0$ & then

$\Delta E_+(R_e, \alpha_{opt}) = -2.35 \text{ eV}$, a significant

improvement.

We could also use the Linear Variation Method to

determine C_A & C_B .

taking $\psi = C_A \psi_A + C_B \psi_B$

& forming

$$E = \frac{\langle \psi | \hat{H} | \psi \rangle}{\langle \psi | \psi \rangle} = E(C_A, C_B)$$

we determine $\frac{\partial E}{\partial C_A} = \frac{\partial E}{\partial C_B} = 0$

which results in the simultaneous equations

$$C_A(H_{AA}-E) + C_B(H_{AB}-ES) = 0$$

$$C_A(H_{AB}-ES) + C_B(H_{BB}-E) = 0$$

The solution requires the secular equation

$$\begin{vmatrix} H_{AA}-E & H_{AB}-ES \\ H_{AB}-ES & H_{BB}-E \end{vmatrix} = 0$$

We have seen $H_{AA} = H_{BB}$, So

$$(H_{AA}-E)^2 = (H_{AB}-ES)^2$$

$$\& \therefore \pm(H_{AA}-E) = (H_{AB}-ES)$$

~~for~~ substituting into the simultaneous equations

$$C_A(H_{AA}-E) + C_B(\pm)(H_{AA}-E) = 0$$

requires $C_A = \pm C_B$

$$\& \therefore \psi_g = C_A(1s_A + 1s_B)$$

$$\psi_u = C_A(1s_A - 1s_B)$$

We can then compute the energy as before or solve for

E in

$$H_{AA}-E = \pm (H_{AB}-ES)$$

Taking the negative sign

$$E_g = \frac{H_{AA} + H_{AB}}{1 + S}$$

we have seen $H_{AA} = -\frac{1}{2} + J$ & $H_{AB} = -\frac{S}{2} + K$

So

$$E_g = \frac{-\frac{1}{2} + J - \frac{S}{2} + K}{1 + S} = -\frac{1}{2} + \frac{J + K}{1 + S} \text{ as found}$$

earlier.

$$\text{Also } E_u = (\text{positive sign}) = \frac{H_{AA} - H_{AB}}{1 - S}$$

$$E_u = \frac{-\frac{1}{2} + J + \frac{S}{2} - K}{1 - S} = -\frac{1}{2} + \frac{J - K}{1 - S}$$

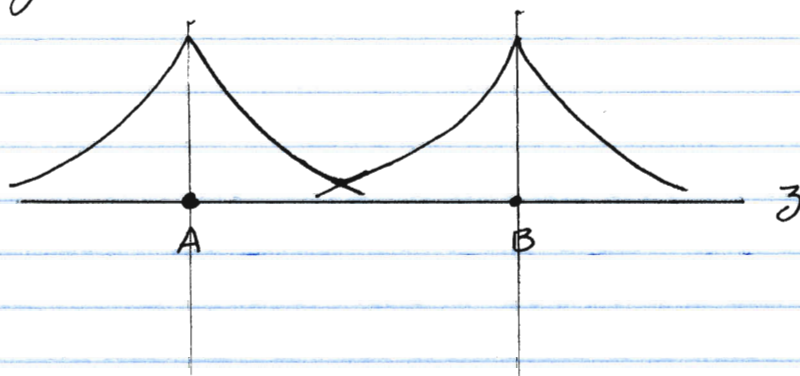
The Ψ_u state does not result in a stable molecule.

Bonding & Anti bonding Orbitals

If one plots the Ψ_g function

$$\Psi_g^2 = \left(\frac{1\psi_A + 1\psi_B}{\sqrt{2 + 2S}} \right)^2$$

along the internuclear line (z) -

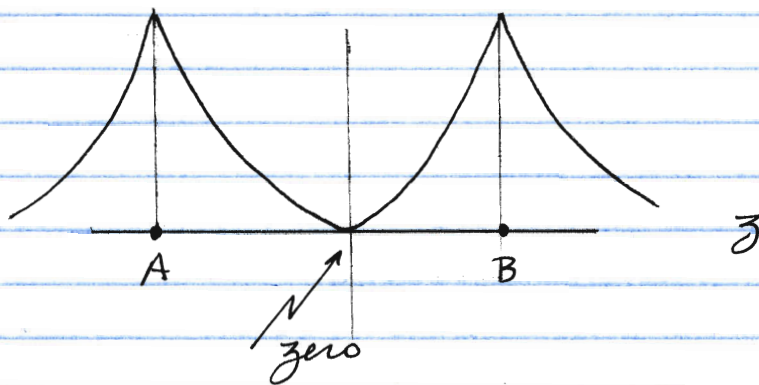


One finds that there is more electron density in the region between the two nuclei than if one simply plotted the sum of the ^{non interacting} atomic densities. The ψ_g is a bonding orbital

On the other hand, the ψ_u function

$$\psi_u = \frac{1s_A - 1s_B}{\sqrt{2}}$$

has a node exactly half way between the two nuclei & when squared the node remains, so ψ_u^2 looks like



& there is a decrease in the electron density relative to the ~~two~~ non interacting atoms. The ψ_u is an antibonding orbital.

Improvements

We took the variation function ^{for the g state} as

$$1A_A + 1A_B$$

but there is no reason to stop with the $1s$ orbitals.

We could, for example write:

$$\begin{aligned} \psi_g = & C_{1s} (1A_A + 1A_B) + C_{2s} (2A_A + 2A_B) + C_{2p} (2P_{3A} - 2P_{3B}) \\ & + C_{3s} (3A_A + 3A_B) + C_{3p} (3P_{3A} - 3P_{3B}) \\ & + C_{3d} (3d_{z^2A} + 3d_{z^2B}) + \dots \end{aligned}$$

* determine the coefficients using the linear Variation method. When we do this we are using the

SCF-LCAO-MO method. The ^{basis} functions

$$1A_A + 1A_B$$

$$2A_A + 2A_B$$

$$2P_{3A} - 2P_{3B}$$

are called symmetry orbitals and one can ^(the C's) write them down immediately. The weight ^{with which}

the symmetry orbitals enter the molecular orbital is determined by the secular equations. The symmetry orbitals can be labelled as

$$\sigma_g 1s = (1s_A + 1s_B) N \rightarrow \text{normalization}$$

$$\sigma_g 2s = (2s_A + 2s_B) N'$$

$$\sigma_g 2p_z = (2p_{zA} + 2p_{zB}) N''$$

$$\sigma_u 1s = (1s_A - 1s_B) N$$

$$\sigma_u 2s = (2s_A - 2s_B) N$$

$$\sigma_u 2p_z = (2p_{zA} - 2p_{zB}) N$$